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52. We thank M. Wagener for excellent assistance with animal experiments and G. Barsh and members of the Kinsley lab for helpful comments on the manuscript. A.M.H. is a trainee of the Medical Scientist Training Program at Stanford, supported by NIH training grant 5T32GM07365 (NIGMS). The initial stages of the genetic cross were supported by a Stanford-SmithKline Beecham research award. D.M.K. is an HHMI assistant investigator.

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# Causes of Climate Change Over the Past 1000 Years

Thomas J. Crowley

Recent reconstructions of Northern Hemisphere temperatures and climate forcing over the past 1000 years allow the warming of the 20th century to be placed within a historical context and various mechanisms of climate change to be tested. Comparisons of observations with simulations from an energy balance climate model indicate that as much as 41 to 64% of preanthropogenic (pre-1850) decadal-scale temperature variations was due to changes in solar irradiance and volcanism. Removal of the forced response from reconstructed temperature time series yields residuals that show similar variability to those of control runs of coupled models, thereby lending support to the models' value as estimates of low-frequency variability in the climate system. Removal of all forcing except greenhouse gases from the ~1000-year time series results in a residual with a very large late-20th-century warming that closely agrees with the response predicted from greenhouse gas forcing. The combination of a unique level of temperature increase in the late 20th century and improved constraints on the role of natural variability provides further evidence that the greenhouse effect has already established itself above the level of natural variability in the climate system. A 21st-century global warming projection far exceeds the natural variability of the past 1000 years and is greater than the best estimate of global temperature change for the last interglacial.

The origin of the late-20th-century increase in global temperatures has prompted considerable discussion. Detailed comparisons of climate model results with observations (1) suggest that anthropogenic changes, particularly greenhouse gas (GHG) increases, are probably responsible for this climate change. However, there are a number of persistent questions with respect to these conclusions that involve uncertainties in the level of low-frequency unforced variability in the climate system (2) and whether factors such as an

increase in solar irradiance or a reduction in volcanism might account for a substantial amount of the observed 20th-century warming (1, 3–10). Although many studies have addressed this issue from the paleoclimate perspective of the past few centuries (3–10), robust conclusions have been hampered by inadequate lengths of the time series being evaluated. Here I show that the agreement between model results and observations for the past 1000 years is sufficiently compelling to allow one to conclude that natural variability plays only a subsidiary role in the 20th-century warming and that the most parsimonious explanation for most of the warming is that it is due to the anthropogenic increase in GHG.

## Data

The data used in this study include physically based reconstructions of Northern Hemisphere temperatures and indices of volcanism, solar variability, and changes in GHGs and tropospheric aerosols.

*Northern Hemisphere temperatures.* Four indices of millennial Northern Hemisphere temperature have been produced over the past 3 years (11–14). The analysis here uses the mean annual temperature reconstructions of Mann *et al.* (11) and of Crowley and Lowery (CL) (12), because the energy balance model used in this study calculates only this term [the other records (13, 14) are estimates of warm-season temperature at mid-high latitudes]. The Mann *et al.* reconstruction was determined (8) by first regressing an empirical orthogonal function analysis of 20th-century mean annual temperatures against various proxy indices (such as tree rings, corals, and ice cores). Past changes in temperature are estimated from variations in the proxy data (15). The Mann *et al.* reconstruction has a varying number of records per unit of time (although the number in the earlier part of the record is still greater than in CL). The CL reconstruction is a more heterogeneous mix of data than the Mann *et al.* reconstruction, but the number of records is nearly constant in time. It is a simple composite of Northern Hemisphere climate records and was scaled (12) to temperature using the instrumental record (16) in the overlap interval 1860–1965. The instrumental record was substituted for the proxy record after 1860 for two reasons: (i) there were too few proxy data in the CL time series after 1965 to reconstruct temperatures for this interval, and (ii) the original CL reconstruction indicated a “warming” over the interval 1885–1925 that is at variance with the instrumental record. This difference has been attributed (11, 17) to

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an early CO<sub>2</sub> fertilization effect on tree growth. The significance of this decision will be further discussed below; model-data correlations presented in the study include both the original proxy record and the substituted instrumental time series.

Despite the different number and types of data and different methods of estimating temperatures, comparison of the decadal smoothed variations in each reconstruction (Fig. 1) indicates good agreement ( $r = 0.73$  for 11-point smoothed correlations over the preanthropogenic interval 1005–1850, with  $P < 0.01$ ). Both records [and the Jones *et al.* (13) and Briffa (14) reconstructions] show the “Medieval Warm Period” in the interval ~1000–1300, a transition interval from about 1300–1580, the 17th-century cold period, the 18th-century recovery, and a cold period in the early 19th century. Even many of the decadal-scale variations in the Medieval Warm Period are reproducible (12), and both reconstructions [and (13, 14)] indicate that peak Northern Hemisphere warmth during the Middle Ages was less than or at most comparable to the mid-20th-century warm period (~1935–1965). This result occurs because Medieval temperature peaks were not synchronous in all records (12). The two temperature reconstructions also agree closely in estimating an ~0.4°C warming between the 17th-century and the mid-20th-century warm period (18).

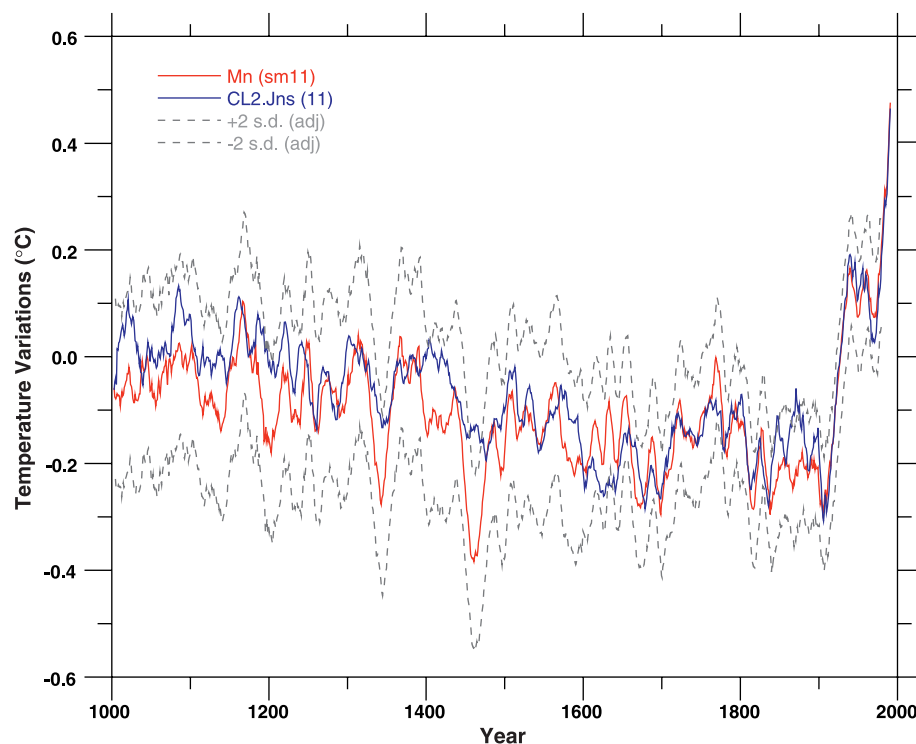
**Volcanic forcing.** There is increasing evidence (3, 7–10) that pulses of volcanism significantly contributed to decadal-scale climate variability in the Little Ice Age. Although some earlier studies (9, 10) of forced climate change back to 1400 used a composite ice core index of volcanism (19), which has a different number of records per unit of time, the present study primarily uses two long ice core records from Crete (20) and the Greenland Ice Sheet Project 2 (GISP2) (21) on Greenland, with a small augmentation from a study of large eruptions recorded in ice cores from both Greenland and Antarctica (22). This approach avoids the potential for biasing model results versus time because of changes in the number of records. Because Southern Hemisphere volcanism north of 20°S influences Northern Hemisphere temperatures, the ice core volcano census samples records down to this latitude. The volcanism record is based on electrical conductivity (20) or sulfate measurements (21), and a catalog of volcanic eruptions (23) was used to remove local eruptions (24) and identify possible candidate eruptions in order to weight the forcing according to latitude. Eruptions of unknown origin were assigned a high-latitude origin unless they also occurred in Antarctic ice

core records (22).

The relative amplitude of volcanic peaks was converted to sulfate concentration by first scaling the peaks to the 1883 Krakatau peak in the ice cores. Although earlier studies (9, 10) linearly converted these concentration changes to radiative forcing changes, subsequent comparison (25) of the very large 1259 eruption [eight times the concentration of sulfate in ice cores from Krakatau and three times the size of the Tambora (1815) eruption (21)] with reconstructed temperatures (11–14) failed to substantiate a response commensurate with a linearly scaled prediction of an enormous perturbation of ~25 W/m<sup>2</sup> (26). Calculations (27) suggest that for stratospheric sulfate loadings greater than about 15 megatons (Mt), increasing the amount of sulfate increases the size of aerosols through coagulation. Because the amount of scattered radiation is proportional to the cross-sectional area, and hence to the 2/3 power of volume (or mass), ice core concentrations estimated as >15 Mt were scaled by this amount (25). Aerosol optical depth was converted to changes in downward shortwave radiative forcing at the tropopause, using the relationship discussed in Sato *et al.* (28). There is significant agreement (29) between the 1000-year-long volcano time series and the concentration-modified Robock and Free (19) times

series (Fig. 2A). Both proxy records show the general trends estimated from ground-based observations of aerosol optical depth (28): the pulse of eruptions in the early 20th century and the nearly 40-year quiescent period of volcanism between about 1920–1960. Because volcano peaks are more difficult to determine in the expanded firn layer of snow/ice cores, updated estimates of Northern Hemisphere radiative forcing from Sato *et al.* were used to extend proxy time series from 1960 to 1998.

**Solar forcing.** There has been much discussion about the effect of solar variability on decadal-to-centennial-scale climates (3, 6, 8–10). An updated version of a reconstruction by Lean *et al.* (5) that spans the interval 1610–1998 was used to evaluate this mechanism [for reference, Free and Robock (10) obtained comparable solar-temperature correlations for the interval 1700–1980 using the Lean *et al.* and alternate Hoyt and Schatten (4) solar reconstructions]. The Lean *et al.* time series has been extended to 1000 by splicing in different estimates of solar variability based on cosmogenic isotopes. These estimates were derived from ice core measurements (30) of <sup>10</sup>Be, residual <sup>14</sup>C from tree ring records (31), and an estimate of <sup>14</sup>C from <sup>10</sup>Be fluctuations (30). The justification for including the latter index is that neither of



**Fig. 1.** Comparison of decadal smoothed Northern Hemisphere mean annual temperature records for the past millennium (1000–1993), based on reconstructions of Mann *et al.* (Mn) (11) and CL (12). The latter record has been spliced into the 11-point smoothed instrumental record (16) in the interval in which they overlap. CL2 refers to a new splice that gives a slightly better fit than the original (12). The autocorrelation of the raw Mann *et al.* time series has been used to adjust (adj) the standard deviation units for the reduction in variance on decadal scales.

the first two splices yields a Medieval solar maximum comparable to that of the present. Because of concerns about biasing results too much by the latter period, which has much more information than the former, the Bard  $^{14}\text{C}$  calculation was included so as to obtain a greater spread of potential solar variations and to allow testing of suggestions (32) that solar irradiance increases could explain the Medieval warming.

Once the splices were obtained, the records were adjusted to yield the potential

$\sim 0.25\%$  change in solar irradiance on longer time scales (33). Because two of the solar proxies indicate that minimum solar activity occurred in the 14th century, the  $0.25\%$  range was set from that time to the present rather than from the 17th century, as was done by Lean *et al.* [the adjustment is very small for the different solar indices in the 14th century ( $\sim 0.05 \text{ W/m}^2$ )]. The 20th-century increase in estimated net radiative forcing from low-frequency solar variability is about 10 to 30% greater than estimated from an independent

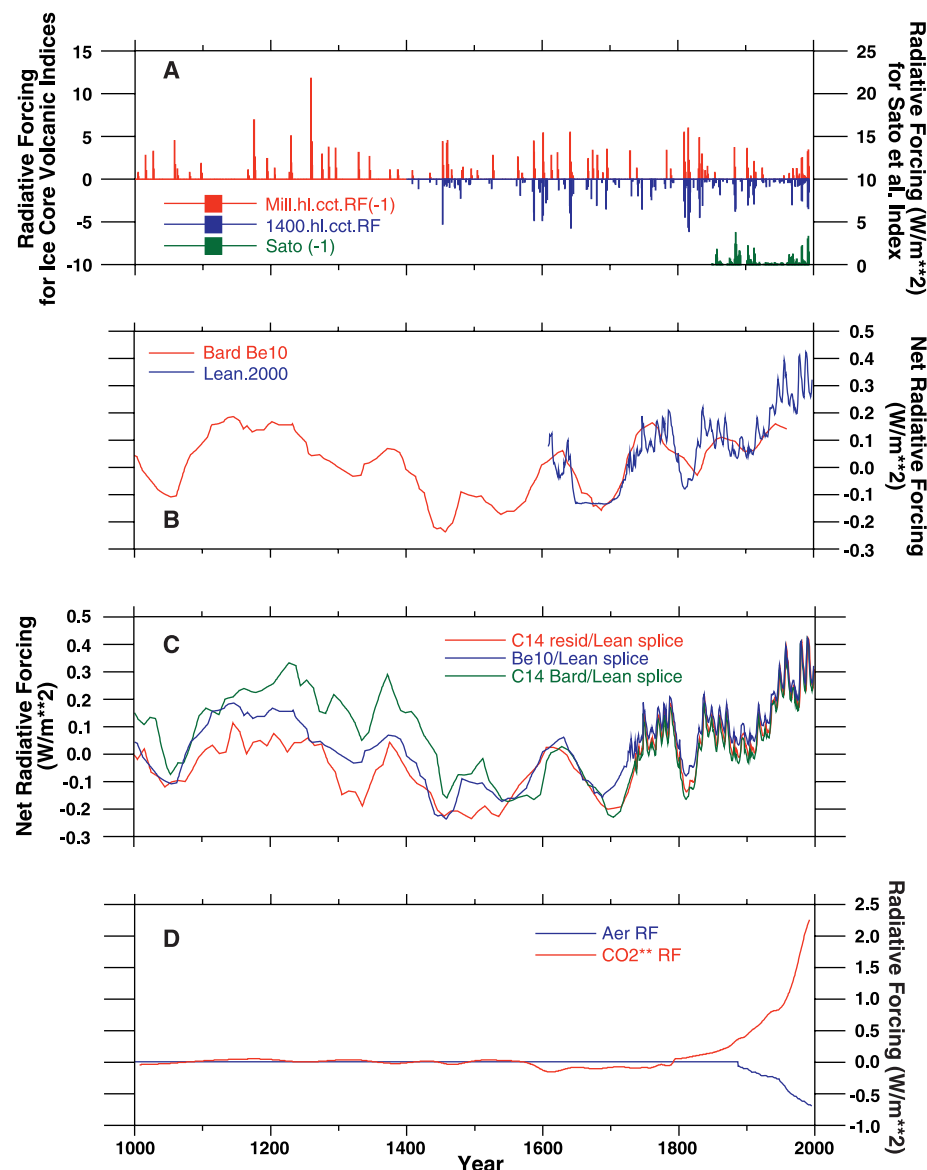
method (34). An example of one of the splices is illustrated in Fig. 2B, and the three composites (Fig. 2C) show the pattern of potential solar variability changes used in this study.

**Anthropogenic forcing.** The standard equivalent radiative forcing for  $\text{CO}_2$  and other well-mixed trace gases (methane, nitrous oxides, and chlorofluorocarbons) is used after 1850 (Fig. 2D). Pre-1850  $\text{CO}_2$  variations, including the small minimum from about 1600–1800, are from Etheridge *et al.* (35). Radiative forcing effects were computed based on updated radiative transfer calculations (36). The well-constrained change in GHG forcing since the middle of the last century is about four times larger than the potential changes in solar variability based on the reconstructions of Lean *et al.* (5) and Lockwood and Stamper (34).

Tropospheric aerosols consider only the direct forcing effect (that is, no cloud feedback), whose global level has been estimated as being about  $-0.4 \text{ W/m}^2$  (37), with the Northern-to-Southern-Hemisphere ratio being in the range of 3 to 4 (38). Because there is an approximate offset in the radiative effects of stratospheric and tropospheric ozone (37), and its total net forcing is on the order of  $+0.2 \text{ W/m}^2$  (37) and is applicable only to the late 20th century, this GHG was not further considered. Other anthropogenic forcing was not included because evaluations by the Intergovernmental Panel for Climate Change (IPCC) (37) indicate that the confidence in these estimates is very low.

## Model

A linear upwelling/diffusion energy balance model (EBM) was used to calculate the mean annual temperature response to estimated forcing changes. This model (39) calculates the temperature of a vertically averaged mixed-layer ocean/atmosphere that is a function of forcing changes and radiative damping. The mixed layer is coupled to the deep ocean with an upwelling/diffusion equation in order to allow for heat storage in the ocean interior. The radiative damping term can be adjusted to embrace the standard range of IPCC sensitivities for a doubling of  $\text{CO}_2$ . The EBM is similar to that used in many IPCC assessments (40) and has been validated (39) against both the Wigley-Raper EBM (40) and two different coupled ocean-atmosphere general circulation model (GCM) simulations (41). All forcings for the model runs were set to an equilibrium sensitivity of  $2^\circ\text{C}$  for a doubling of  $\text{CO}_2$ . This is on the lower end of the IPCC range (42) of  $1.5^\circ$  to  $4.5^\circ\text{C}$  for a doubling of  $\text{CO}_2$  and is slightly less than the IPCC “best guess” sensitivity of  $2.5^\circ\text{C}$  [the inclusion of solar variability in model calculations can decrease the best fit sensitivity (9)]. For both the solar and volcanism runs,



**Fig. 2.** Forcing time series used in the model runs (note scale changes for different panels). (A) (Red) Ice core millennial volcanism time series from this study (multiplied by  $-1$  for display purposes); (blue) ice-core Robock and Free (79) reconstruction from 1400 to the present after adjustments discussed in (9) and (25); and (green) Sato *et al.* (28) Northern Hemisphere radiative forcing, updated to 1998 and multiplied by  $-1$  for display purposes. (B) Example of splice for solar variability reconstructions, using the  $^{10}\text{Be}$ -based irradiance reconstruction of (30) (red) and the reconstruction of solar variability from Lean *et al.* (5) (blue). (C) Comparison of three different reconstructions of solar variability based on  $^{10}\text{Be}$  measurements (30) (blue),  $^{14}\text{C}$  residuals (31) (red), and calculated  $^{14}\text{C}$  changes based on  $^{10}\text{Be}$  variations (30) (green). (D) Splice of  $\text{CO}_2$  radiative forcing changes 1000–1850 (35) (red) and post-1850 anthropogenic changes in equivalent GHG forcing and tropospheric aerosols (blue).



the calculated temperature response is based on net radiative forcing after adjusting for the 30% albedo of the Earth-atmosphere system over visible wavelengths.

## Results

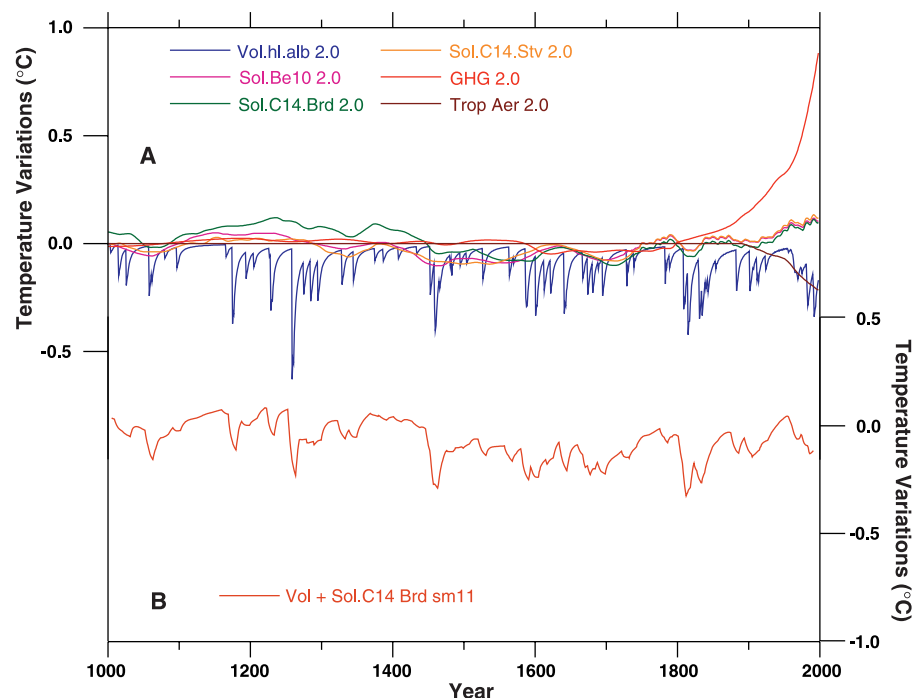
The modeled responses to individual forcing terms (Fig. 3A) indicate that the post-1850 GHG and tropospheric aerosol changes are similar to those discussed in IPCC (42). CO<sub>2</sub> temperature variations are very small for the preanthropogenic interval, although there is a 0.05°C decrease in the 17th and 18th centuries that reflects the CO<sub>2</sub> decrease of ~6 parts per million in the original ice core record (35). Solar variations are on the order of 0.2°C, and volcanism causes large cooling (43) in the Little Ice Age (3–7, 9, 10). Averaged over the entire preanthropogenic interval (Table 1), 22 to 23% of the decadal-scale variance can be explained by volcanism ( $P < 0.01$ ). However, over the interval 1400–1850, the volcanic contribution increases to 41 to 49% ( $P < 0.01$ ), thereby indicating a very important role for volcanism during the Little Ice Age.

The sun-climate correlations for the interval 1000–1850 vary substantially by choice of solar index (Table 1), with explained variance ranging from as low as 9% ( $P < 0.01$ ) for the <sup>14</sup>C residual index (31) to as high as 45% ( $P < 0.01$ ) for the Bard *et al.* (30) <sup>14</sup>C solar index, which reconstructs a Medieval solar warming comparable to the present century but only about 0.1°C greater than predicted by the other solar indices (Fig. 3A). The large range in correlations for the solar records emphasizes the need to determine more precisely the relative magnitude of the real Medieval solar warming peak.

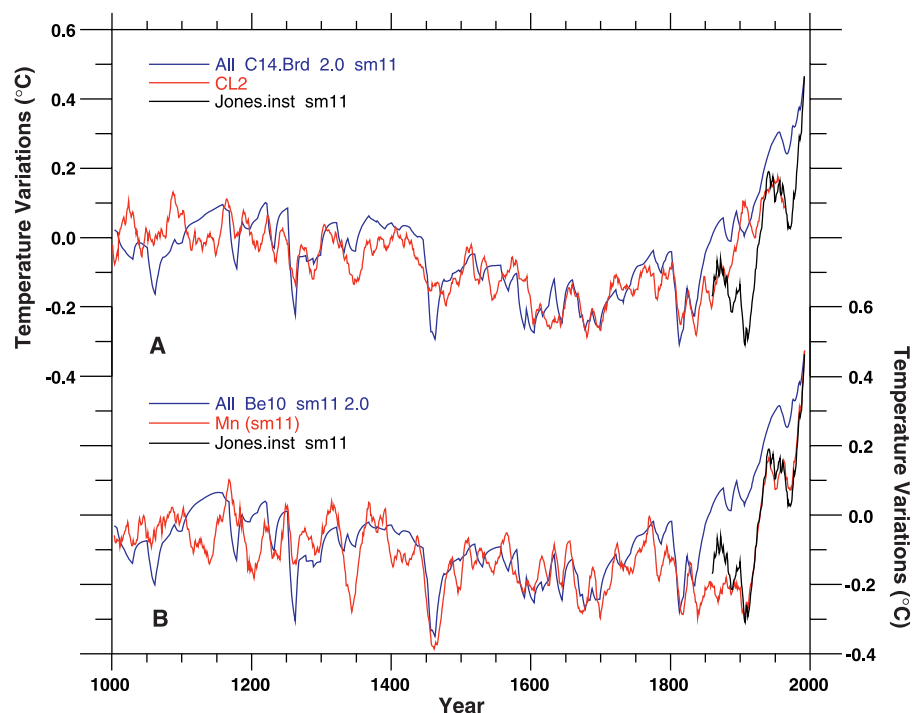
The joint effects of solar variability and volcanism (Fig. 3B) indicate that the combination of these effects could have contributed

0.15° to 0.2°C to the temperature increase (Fig. 1) from about 1905–1955, but only about one-quarter to the total 20th-century

warming. The combined warmth produced by solar variability and volcanism in the 1950s is similar in magnitude but shorter in duration



**Fig. 3.** (A) Model response to different forcings, calculated at a sensitivity of 2.0°C for a doubling of CO<sub>2</sub>. (B) Example of the combined effect of volcanism and solar variability (with 11-point smoothing), using the Bard *et al.* (30) <sup>14</sup>C index.



**Fig. 4.** Comparison of model response (blue) using all forcing terms (with a sensitivity of 2.0°C) against (A) the CL (12) data set spliced into the 11-point smoothed Jones *et al.* (16) Northern Hemisphere instrumental record, with rescaling as discussed in the text and in the Fig. 1 caption; and (B) the smoothed Mann *et al.* (11) reconstruction. Both panels include the Jones *et al.* instrumental record for reference. To illustrate variations in the modeled response, the <sup>14</sup>C calculation from Bard *et al.* (30) has been used in (A) and the <sup>10</sup>Be estimates from (30) have been used in (B).

**Table 1.** Correlations of volcanism (volc.) and solar variability (sol.) for the preanthropogenic interval, with percent variance shown in parentheses. The different solar time series reflect the three different solar indices used in this study. The Mann *et al.* time series (17) has been smoothed with an 11-point filter. CL was smoothed in the original analysis (12). Different abbreviations for solar forcing refer to the different indices discussed in the text: <sup>10</sup>Be and <sup>14</sup>C calculations are from Bard *et al.* (30); <sup>14</sup>C residuals are from Stuiver *et al.* (31).

|                                                       |            |
|-------------------------------------------------------|------------|
| Volc. vs. Mann <i>et al.</i> (1000–1850)              | 0.48 (23%) |
| Volc. vs. CL (1000–1850)                              | 0.47 (22%) |
| Volc. vs. Mann <i>et al.</i> (1400–1850)              | 0.70 (49%) |
| Volc. vs. CL (1400–1850)                              | 0.64 (41%) |
| Sol ( <sup>10</sup> Be) vs. Mann <i>et al.</i>        | 0.45 (20%) |
| Sol ( <sup>14</sup> C Bard) vs. Mann <i>et al.</i>    | 0.56 (31%) |
| Sol ( <sup>14</sup> C Stuiver) vs. Mann <i>et al.</i> | 0.37 (14%) |
| Sol ( <sup>10</sup> Be) vs. CL                        | 0.42 (18%) |
| Sol ( <sup>14</sup> C Bard) vs. CL                    | 0.67 (45%) |
| Sol ( <sup>14</sup> C Stuiver) vs. CL                 | 0.30 (9%)  |

than the warmth simulated by these mechanisms in the Middle Ages. The variations in the past few decades resulting from the com-

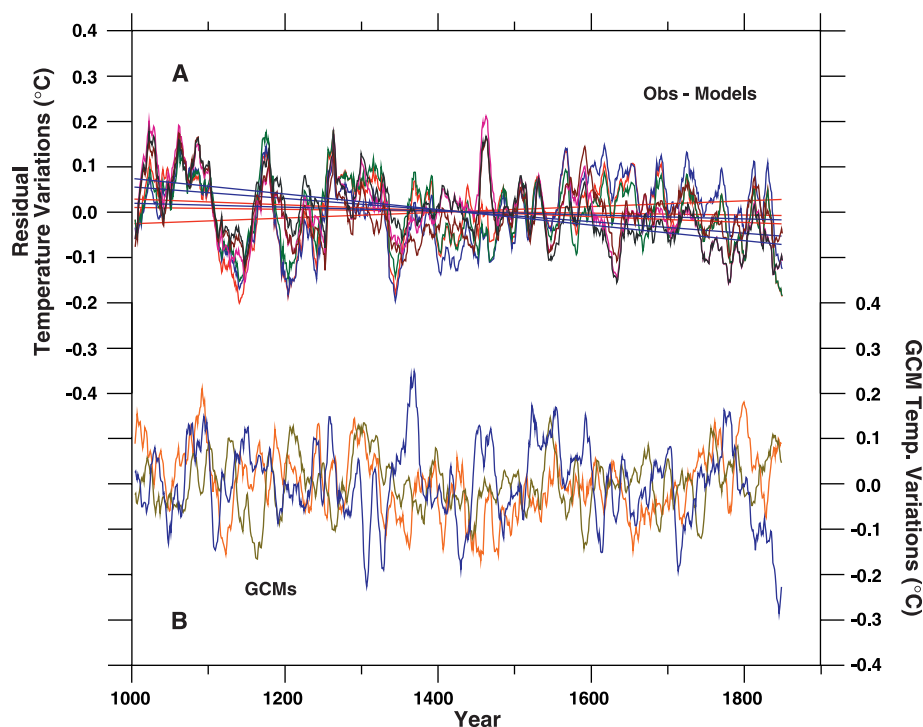
bination of solar variability and volcanism is 0.2°C less than the 1955 peak.

Combining all forcing (solar, volcanism,

GHG, and tropospheric aerosols) results in some striking correspondences between the model and the data over the preanthropogenic interval (Fig. 4). Eleven-point smoothed correlations (44) for the preanthropogenic interval (Table 2) indicate that 41 to 64% of the total variance is forced ( $P < 0.01$ ). The highest correlations are obtained for the CL time series, which has slightly more Medieval warmth than the Mann *et al.* reconstruction, and for the forcing time series that includes the largest solar estimate of Medieval warmth. Forced variability explains 41 to 59% of the variance ( $P < 0.01$ ) over the entire length of the records. Although simulated temperatures agree with observations in the late 20th century, simulations exceed observations by  $\sim 0.1^\circ$  to  $0.15^\circ\text{C}$  over the intervals 1850–1885 and 1925–1975, with a larger discrepancy between  $\sim 1885$ –1925 that reaches a maximum offset of  $\sim 0.3^\circ\text{C}$  from  $\sim 1900$ –1920. However, decadal-scale patterns of warming and cooling are still simulated well in these offset intervals. A sensitivity test (45) comparing forcing time series with and without solar variability indicates that changes caused by volcanism and  $\text{CO}_2$  are responsible for the simulated temperature increase from the mid- to late 19th century to the early 20th century, thereby eliminating uncertainties in solar forcing as the explanation for the temperature differences between the model and the data. Also shown in Fig. 4A is the CL reconstruction with the “anomalous” warm interval ( $\sim 1885$ –1925) discussed above. For this reconstruction, 55 to 69% of the variance from 1005–1993 can be explained by the model ( $P < 0.01$ ).

Another means of evaluating the role of forced variability is to determine residuals by subtracting the different model time series from the two paleo time series over the preanthropogenic interval (Fig. 5A). The trend lines for three of these residuals are virtually zero, and there is only about a  $\pm 0.1^\circ\text{C}$  trend for the other three residuals. Because the pre-1850 residuals represent an estimate of the unforced variability in the climate system, it is of interest to compare the smoothed residuals with smoothed estimates of unforced variability in the climate system from control runs of coupled ocean-atmosphere models. There is significant agreement (Fig. 5B and Table 3) between the smoothed standard deviations of the GCMs (46) and paleo residuals (47). These results support a basic assumption in optimal detection studies (1) and previous conclusions (48) that the late-20th-century warming cannot be explained by unforced variability in the ocean-atmosphere system. However, a combination of GHG, natural forcing, and ocean-atmosphere variability could have contributed to the 1930–1960 warm period (1, 9, 10, 49).

One way to highlight the unusual nature



**Fig. 5.** Analysis of preanthropogenic residuals in the paleo records. **(A)** Estimates of residuals using all combinations of temperature reconstructions and total forcing (including three different solar indices), with trend lines fitted for each of the six residuals. **(B)** Control runs (detrended) from three different coupled ocean-atmosphere models (46): the NOAA/GFDL model (NOAA Geophysical Fluid Dynamics Laboratory) (orange), the HadCM3 model (Hadley Centre at the UK Meteorological Office, Bracknell, UK) (blue), and the ECHAM3/LSG model (European Centre/University of Hamburg/Max Planck Institute für Meteorologie) (brown). For the sake of comparison with the paleo data, the GCM runs have been truncated to the same length as the paleo residuals and have been plotted using the arbitrary starting year of 1000.

**Table 2.** Correlations between model runs with combined forcing and the Mann *et al.* (11) and CL (12) time series. Correlations have been subdivided into the following three categories: Top set: Correlations for all the preanthropogenic interval 1005–1850 of model response to combined forcing (“All”) with different solar indices (Table 1) and the 11-point smoothed Mann *et al.* time series and CL2 record spliced into the 11-point smoothed Jones *et al.* (16) time series. Middle set: Correlations over the entire interval analyzed. Bottom set: Correlations and variance explained for the interval 1005–1993 using the original CL2 reconstruction from 1005–1965, with the smoothed Jones *et al.* (16) record added from 1965–1993.

| Summary of pre-1850 correlations, with variance shown in parentheses                                                                                      |            |
|-----------------------------------------------------------------------------------------------------------------------------------------------------------|------------|
| All $^{10}\text{Be}$ (solar) vs. Mann (sm11)                                                                                                              | 0.64 (41%) |
| All $^{14}\text{C}$ Brd (solar) vs. Mann (sm11)                                                                                                           | 0.68 (46%) |
| All $^{14}\text{C}$ Stv (solar) vs. Mann (sm11)                                                                                                           | 0.65 (42%) |
| All $^{10}\text{Be}$ (solar) vs. CL2.Jns11                                                                                                                | 0.69 (48%) |
| All $^{14}\text{C}$ Brd (solar) vs. CL2.Jns11                                                                                                             | 0.80 (64%) |
| All $^{14}\text{C}$ Stv (solar) vs. CL2.Jns11                                                                                                             | 0.68 (47%) |
| Summary of correlations for 1005–1993, with variance shown in parentheses                                                                                 |            |
| All $^{10}\text{Be}$ (solar) vs. Mann (sm11)                                                                                                              | 0.68 (46%) |
| All $^{14}\text{C}$ Brd (solar) vs. Mann (sm11)                                                                                                           | 0.73 (53%) |
| All $^{14}\text{C}$ Stv (solar) vs. Mann (sm11)                                                                                                           | 0.67 (45%) |
| All $^{10}\text{Be}$ (solar) vs. CL2.Jns11                                                                                                                | 0.66 (43%) |
| All $^{14}\text{C}$ Brd (solar) vs. CL2.Jns11                                                                                                             | 0.77 (59%) |
| All $^{14}\text{C}$ Stv (solar) vs. CL2.Jns11                                                                                                             | 0.64 (41%) |
| Summary of correlations for 1005–1993 against unfiltered CL time series, with 11-point smoothed Jones <i>et al.</i> (16) record spliced in from 1965–1993 |            |
| All $^{10}\text{Be}$ (solar) vs. CL2.Jns 11                                                                                                               | 0.75 (57%) |
| All $^{14}\text{C}$ Brd (solar) vs. CL2.Jns11                                                                                                             | 0.83 (69%) |
| All $^{14}\text{C}$ Stv (solar) vs. CL2.Jns11                                                                                                             | 0.74 (54%) |

of the late-20th-century warmth is to subtract all forcing other than CO<sub>2</sub> (solar, volcanism, and tropospheric aerosols) and examine the late-20th-century residuals within the context of the previous 1000 years (Fig. 6). There is an unprecedented residual warming in the late 20th century that matches the warming predicted by GHG forcing. Projection of the "Business As Usual" (BAU) scenario into the next century using the same model sensitivity of 2.0°C indicates that, when placed in the perspective of the past 1000 years, the warming will reach truly extraordinary levels (Fig. 6). The temperature estimates for 2100 also exceed the most comprehensive estimates (50) of global temperature change during the last interglacial (~120,000 to 130,000 years ago)—the warmest interval in the past 400,000 years.

## Discussion

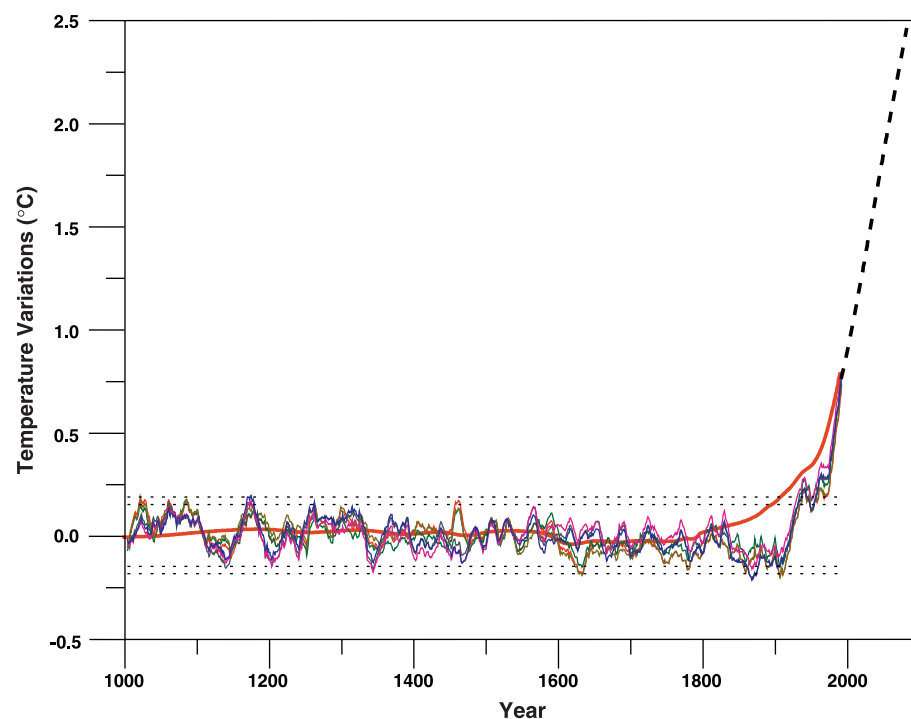
Forcing a linear energy balance model with independently derived time series of volcanism and solar variability indicates that 41 to 64% of the preanthropogenic low-frequency variance in temperature can be accounted for by external factors. These results were obtained without any retuning of the climate model. When the same sensitivity for the preanthropogenic interval is used for the past 150 years, there is good agreement with temperatures in the late 20th century. Some caution is needed in interpreting the agreement between models and data for a 2.0°C sensitivity, because a more detailed analysis of uncertainties (47) might yield slightly different sensitivities than simulated here (51). Also, statistical methods better constrain the minimum than the maximum sensitivity (52). If paleo records are shown to have had larger amplitude than used in this study (18), model-data correlations should still be valid but the best fit sensitivity would be greater.

The largest model-data discrepancy over the entire past 1000 years is from ~1885–1925, peaking in ~1900–1920. Although such differences could reflect random uncertainties in the paleo reconstructions (Fig. 1) or forcing fields, the consistent offset between model and data suggests the need to identify one or more specific explanations for the differences. For example, two factors that could be contributing to the model-data differences in this interval are: (i) mid-latitude land clearance may have increased albedo and caused slightly greater cooling than simulated (53), and (ii) warming may be underestimated in the early stage of the instrumental record because of sparse data coverage (16). As discussed above, there is evidence for warming in some of the high-elevation data in the original CL (12) reconstruction and in the comprehensive Overpeck *et al.* (6) Arctic synthesis. Many alpine glaciers started to retreat around 1850 (54). There is also

some evidence for warming of tropical oceans in the late 19th century (16), but the data are very sparse. More data and model analyses would be required to test these and other possible explanations.

Analysis of residuals in the pre-1850 interval reveals little or no trend. If the pre-

1850 temperature reconstructions, forcing estimates, and model responses are correct, the model-data agreement in this study suggests that factors such as thermohaline circulation changes (55) may have played only a secondary role with respect to modifying hemispheric temperatures over the past 1000 years (56).



**Fig. 6.** Comparison of the GHG forcing response (from Fig. 3) with six residuals determined by removing all forcing except GHG from the two different temperature reconstructions in Fig. 1. As in Fig. 5, the three different estimates of solar variability were used to get one estimate of the uncertainty in the response. This figure illustrates that GHG changes can explain the 20th-century rise in the residuals;  $\pm 2$  standard deviation lines (horizontal dashed lines) refer to maximum variability of residuals from Fig. 5A (inner dashes) and maximum variability (outer dashes) of the original pre-1850 time series (Fig. 1). The projected 21st-century temperature increase (heavy dashed line at right) uses the IPCC BAU scenario (the "so-called IS92a forcing") for both GHG and aerosols (sulfate and biomass burning, including indirect effects), and the model simulation was run at the same sensitivity (2.0°C for a doubling of CO<sub>2</sub>) as other model simulations in this article. The IS92a scenario is from (59).

**Table 3.** Comparison of smoothed standard deviations (in °C) of 850-year Northern Hemisphere preanthropogenic residuals from observations with smoothed 850-year coupled ocean-atmosphere GCM control runs (46). All records were detrended except for the original smoothed paleo time series.

|                                                 |                  |
|-------------------------------------------------|------------------|
| Mann <i>et al.</i> (sm11)                       | 0.066°–0.075°C   |
| CL                                              | 0.054°–0.061°C   |
| ECHAM3/LSG* (sm11)                              | 0.058°C          |
| GFDL† (sm11)                                    | 0.072°C          |
| HadCM3‡ (sm11)                                  | 0.086°C          |
| Summary statistics for 11-year smoothing        |                  |
| Paleo residuals                                 | 0.064° ± 0.009°C |
| GCMs                                            | 0.072° ± 0.014°C |
| Summary statistics for 51-year smoothing        |                  |
| Paleo residuals                                 | 0.042° ± 0.010°C |
| GCMs                                            | 0.044° ± 0.010°C |
| Original smoothed paleo time series (1005–1850) |                  |
| Mann <i>et al.</i>                              | 0.086°C          |
| CL                                              | 0.092°C          |

\*The European Centre/University of Hamburg/Max Planck Institute für Meteorologie model. †The NOAA Geophysical Fluid Dynamics Laboratory model, Princeton, New Jersey. ‡The Hadley Centre model at the UK Meteorological Office, Bracknell, UK.



Mann *et al.* (11) suggested that the decrease in summer insolation since the early Holocene (57) also could have contributed to the cooling between the Middle Ages and the Little Ice Age. However, calculations (58) do not support this suggestion.

There are therefore two independent lines of evidence pointing to the unusual nature of late-20th-century temperatures. First, the warming over the past century is unprecedented in the past 1000 years. Second, the same climate model that can successfully explain much of the variability in Northern Hemisphere temperature over the interval 1000–1850 indicates that only about 25% of the 20th-century temperature increase can be attributed to natural variability. The bulk of the 20th-century warming is consistent with that predicted from GHG increases. These twin lines of evidence provide further support for the idea that the greenhouse effect is already here. This assertion may seem surprising to some because of continuing uncertainties with respect to the dynamical response of the ocean-atmosphere system and radiative forcing feedbacks (both direct and indirect) of, for example, clouds, biomass burning, and mineral dust. Although regional climate change is almost certainly influenced by these complex dynamic and thermodynamic feedbacks, the striking agreement seen in this study between simple model calculations and observations indicates that on the largest scale, temperature responds almost linearly to the estimated changes in radiative forcing. The very good agreement between models and data in the preanthropogenic interval also enhances confidence in the overall ability of climate models to simulate temperature variability on the largest scales.

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## REPORTS

# Signature of Superfluid Density in the Single-Particle Excitation Spectrum of $\text{Bi}_2\text{Sr}_2\text{CaCu}_2\text{O}_{8+\delta}$

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We report that the doping and temperature dependence of photoemission spectra near the Brillouin zone boundary of  $\text{Bi}_2\text{Sr}_2\text{CaCu}_2\text{O}_{8+\delta}$  exhibit unexpected sensitivity to the superfluid density. In the superconducting state, the photoemission peak intensity as a function of doping scales with the superfluid density and the condensation energy. As a function of temperature, the peak intensity shows an abrupt behavior near the superconducting phase transition temperature where phase coherence sets in, rather than near the temperature where the gap opens. This anomalous manifestation of collective effects in single-particle spectroscopy raises important questions concerning the mechanism of high-temperature superconductivity.

The collective nature of superconductivity manifests itself contrastingly in different techniques. Microwave and muon spin relaxation measurements are inherently sensitive to the collective motion of the condensate, whereas single-electron tunneling spectroscopy and photoemission mainly probe single-particle excitations of the condensate. Hence, these two types of spectroscopies can be used to measure two essential but distinct ingredients of superconductivity: the

superfluid density, which characterizes the phase coherence of the Cooper pairs, and the superconducting energy gap, which reflects the strength of the pairing. We report a pronounced departure from this conventional picture on the basis of angle-resolved photoemission spectroscopy (ARPES) data from  $\text{Bi}_2\text{Sr}_2\text{CaCu}_2\text{O}_{8+\delta}$  (Bi2212).

In Bi2212, a well-known peak and dip feature develops near the superconducting phase transition temperature  $T_c$  (1–6) in the ARPES spectra near the Brillouin zone boundary ( $\pi/a, 0$ ), where  $a$  is the lattice constant, which is set to unity for convenience. This feature has so far been discussed primarily in the context of quasi-particle excitations coupled to many-body collective excitations (3, 7, 8). Here, we show that the doping dependence of the peak intensity exhibits a clear resemblance to the behavior exhibited by the superfluid density  $n_s$  and the conden-

sation energy, both of which scale approximately with dopant  $x$  in the underdoped regime and saturate or even scale with  $A - x$  in the overdoped regime (where  $A$  is a constant). The temperature dependence of this peak intensity also shows a resemblance to that of the superfluid density. More important, this peak intensity shows an abrupt behavior near  $T_c$ , where phase coherence sets in, rather than at  $T^*$ , the temperature where the pseudogap opens in the underdoped regime (9). It is remarkable that the signature of these collective properties appears in a single-particle excitation spectrum at  $(\pi, 0)$  (the antinode region of a d-wave state with maximum gap).

This anomalous manifestation of the superfluidity as well as  $x$  dependence of many physical quantities contrasts strongly with the conventional Bardeen-Cooper-Schrieffer (BCS) type of picture based on the Fermi liquid. In that picture, the quasi-particle spectral weight  $Z$  depends on interactions and the energy gap near the normal state Fermi surface whose volume scales with  $1 - x$  (counting electrons), rather than on the superfluid density. Instead, these observations agree well with theories that are based on the doped Mott insulator.

We measured ARPES spectra on Bi2212 samples with various doping levels (10). Bi2212 samples are labeled by their  $T_c$  with the prefix UD for underdoped, OP for optimally doped, or OD for overdoped (e.g., an underdoped  $T_c = 83$  K sample is denoted UD83). Samples used here include traveling-solvent floating zone-grown single crystals and molecular beam epitaxy (MBE)-grown films. The typical transition width is less than 1 K, except for UD73, which has a transition width of 7 K. Samples with different  $T_c$ 's are of similar high quality, as assessed by the measured residual resistivity ratio (RRR), the ratio between the extrapolated resistivity at  $T = 0$  K and resistivity at  $T = 300$  K. The hole doping level  $x$  was determined by the empirical relation  $T_c = T_{c,\text{max}}[1 - 82.6(x - 0.16)^2]$  (11).  $T_{c,\text{max}} = 91$  K

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